

Experiment and ImQMD model comparison using $^{129,124}\text{Xe} + ^{124,112}\text{Sn}$ @ 100AMeV collisions

**Korea-China joint workshop
for rare isotope physics**



Korea University
Seon ho Nam
2025. 7. 10. Thu

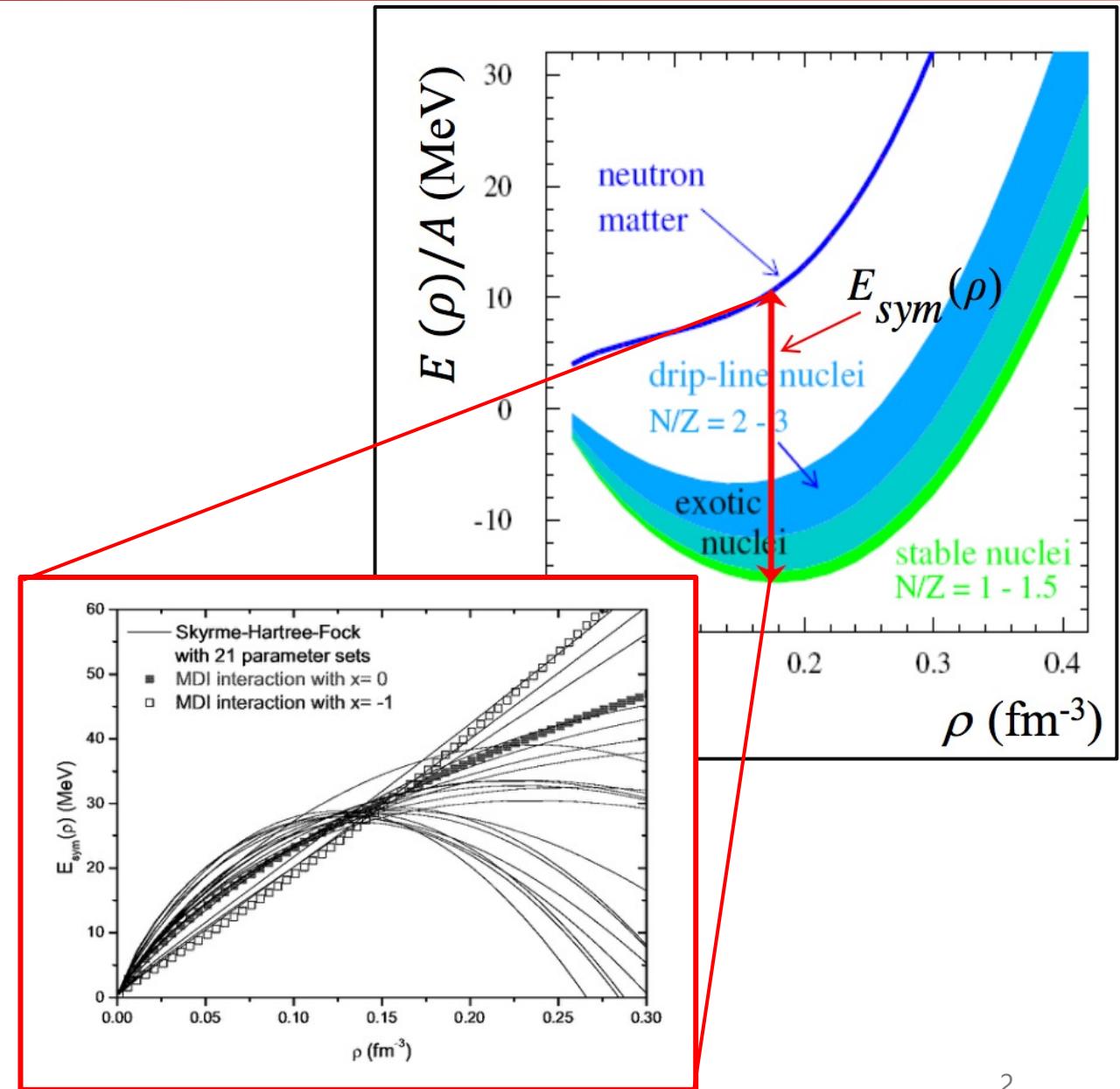
Nuclear symmetry energy

Nuclear Equation of State(NEoS)

$$E(\rho, \delta) = E(\rho, 0) + E_{sym}(\rho)\delta^2 + O(\delta^4)$$

$$\text{where, } \delta = \frac{\rho_n - \rho_p}{\rho}$$

$$\begin{aligned} E_{sym}(\rho) &= E_{sym}(\rho_0) \left(\frac{\rho}{\rho_0}\right)^{\gamma} \\ &= S + \frac{L_{sym}}{3} \left(\frac{\rho - \rho_0}{\rho_0}\right) + \frac{K_{sym}}{18} \left(\frac{\rho - \rho_0}{\rho_0}\right)^2 \dots \end{aligned}$$



R_{n/p} Slope

(PHYSICAL REVIEW C 109, 054624 (2024))

Thermodynamic picture of primary fragment yield

$$Y_{pri}(N, Z, T) \approx \frac{VA^{3/2}}{\lambda_T^3} \omega(N, Z, T) \exp\left[\frac{E_B(N, Z) + N\mu_n + Z\mu_p}{T}\right]$$

$$Y(N, Z, T) = f(N, Z, T) Y_{pri}(N, Z, T)$$

$$1) R_{n/p} = \frac{Y(n)}{Y(p)} \propto \exp\left(\frac{\mu_n - \mu_p}{T}\right) = \exp\left(\frac{2V_{sym}\delta}{T}\right)$$

$$2) V_{sym}(p) = V_{sym}^0 + \frac{\partial V_{sym}}{\partial p^2} p^2 \dots$$

$$3) \Delta m_{np}^* \approx -\left(\frac{m^*}{m}\right)^2 4m\delta \frac{\partial V_{sym}}{\partial p^2}$$

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1) + 2) + 3) →

$$R_{n/p} \approx \exp\left(\frac{2V_{sym}^0\delta}{T}\right) \exp\left(-\frac{\lambda\left(\frac{m^*}{m}\right)^2 \Delta m_{np}^*}{T} (E_k/A)\right)$$

$$\textcolor{red}{S_{n/p}} = \frac{\partial \ln R_{n/p}}{\partial E_k/A} \propto -\frac{\lambda}{T} \left(\frac{m}{m^*}\right)^2 \Delta m_{np}^*$$

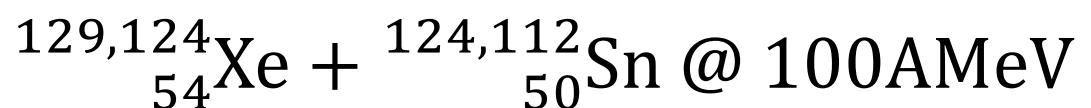
$$\mathbf{DR(n/p)} = \mathbf{R_{n/p}^{higher\ N/Z}} / \mathbf{R_{n/p}^{lower\ N/Z}}$$

$$Y_{C.I.\textcolor{blue}{n}(\textcolor{red}{p})} = \frac{dM_{\textcolor{blue}{n}(\textcolor{red}{p})\ C.I.}}{dE_{c.m.} dy_{c.m.}}$$

$$= \sum_{N,Z} \textcolor{blue}{N}(\textcolor{red}{Z}) \frac{dM(N, Z)}{d(E/A)_{c.m.} dy_{c.m.}}$$

Analysis condition and Experiment

- INDRA 4th campaign at GSI (1998-1999)
(INDRA-ALADIN collaboration)
- Bayesian estimation method



	IW1
$e t_{12} [\text{MeV}]$	750~1050
b_0	0.21~0.42

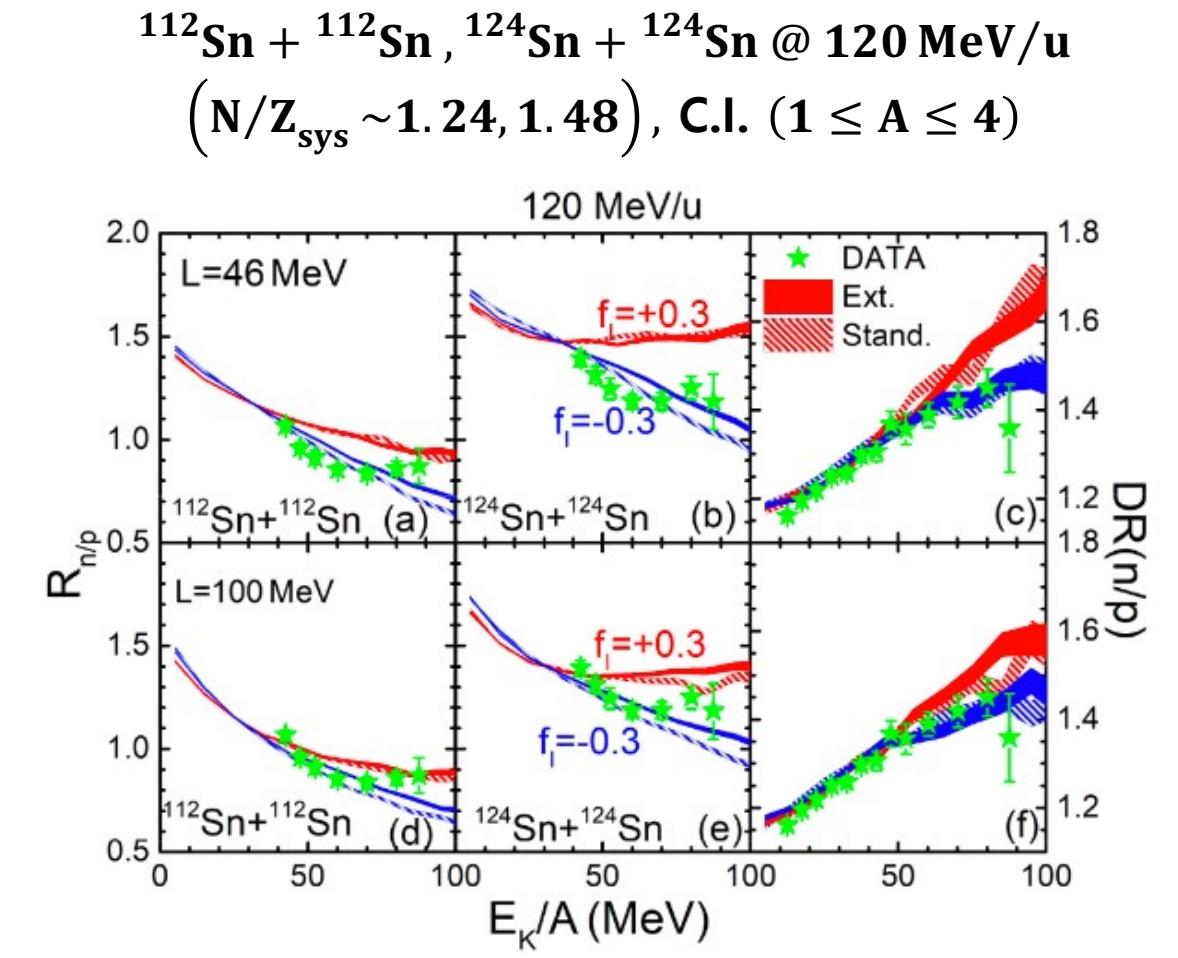
- No neutron detection :
C.I. using $2 \leq A \leq 4$

- Rapidity selection :
 $-0.4 < y_0^{\text{cm}} < 0.4$

ImQMD(+ GEMINI++) with Skyrme parameter

Para	ρ_0	E_0	K_0	S_0	L
SLy4	0.160	-15.97	230	32	46
SkM*	0.160	-15.77	217	30	46
	K_{sym}	m^*/m	m_n^*/m	m_p^*/m	
SLy4	-120	0.69	0.68	0.71	
SkM*	-156	0.79	0.82	0.76	

- 5 discrete b_0 are chosen for time efficiency.
(0.1, 0.2, 0.3, 0.4 and 0.5) \approx IW1
- About 10^6 events for each b_0



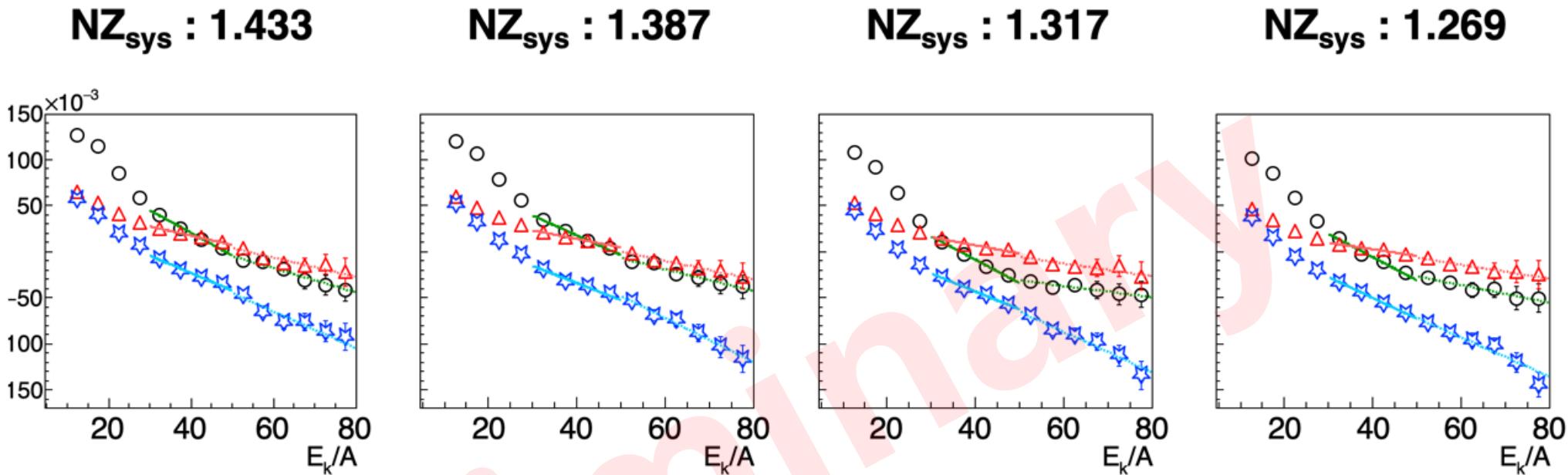
Red($f_I: 0.3$) : $m_n^* < m_p^*$ Blue($f_I: -0.3$) : $m_n^* > m_p^*$

(PHYSICAL REVIEW C 109, 054624 (2024))

Analysis results

$^{129,124}\text{Xe} + ^{124,112}\text{Sn}$ @ 100 MeV/u

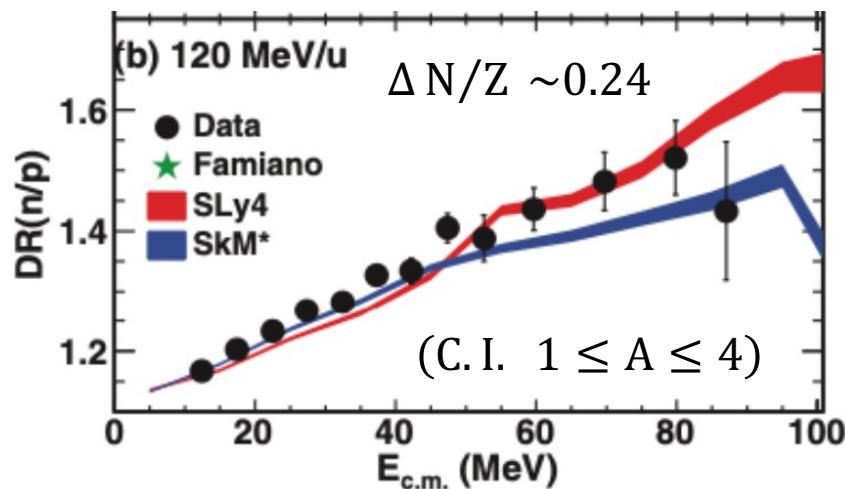
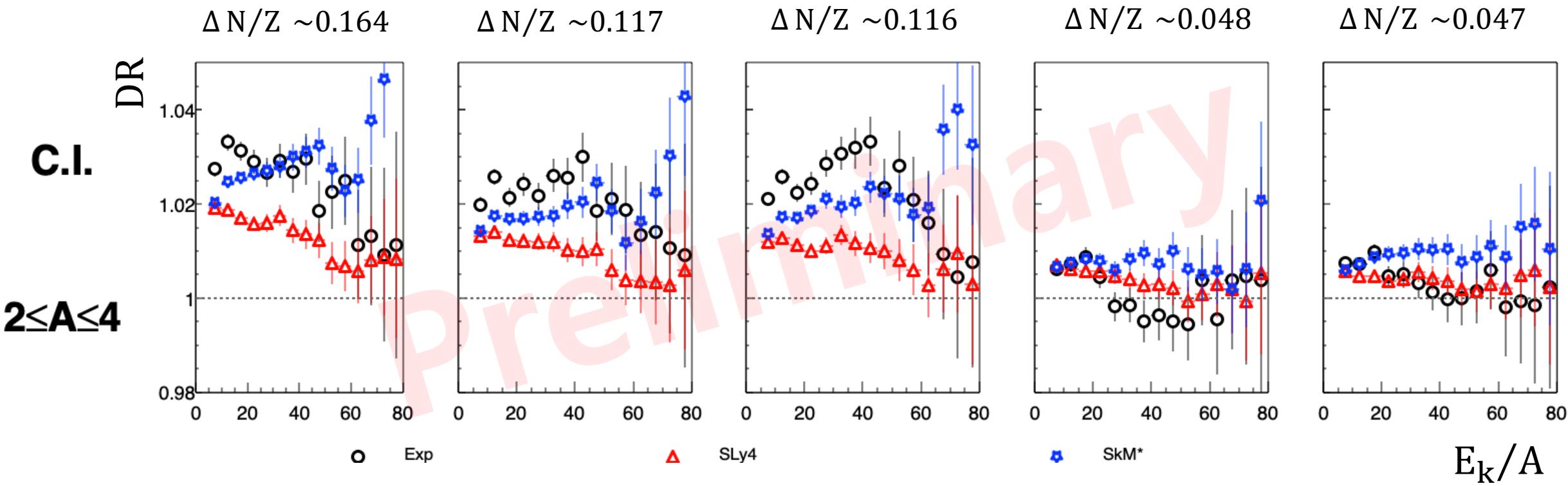
$\ln R_{n/p}$ C.I.
 $2 \leq A \leq 4$



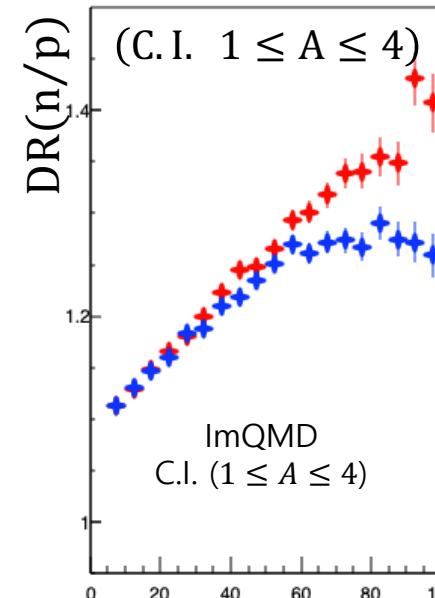
○ Exp

* SkM*

△ SLy4



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- Mass splitting studies using $S_{n/p}$ of yield ratio is consistent with previous studies.
- Different energy dependency and relatively scale of two parameter sets are related to absence of free nucleons.

Summary

1. To study symmetry energy, $S_{n/p}$ and DR(n/p) from experimental and ImQMD calculation are compared.
2. Like previous study, $S_{n/p}$ well reflects effective mass splitting even free nucleons are excluded.
3. DR(n/p) seems less sensitive to effective mass splitting than $S_{n/p}$.
4. Discussion of the analysis results is currently ongoing.

Thank you!

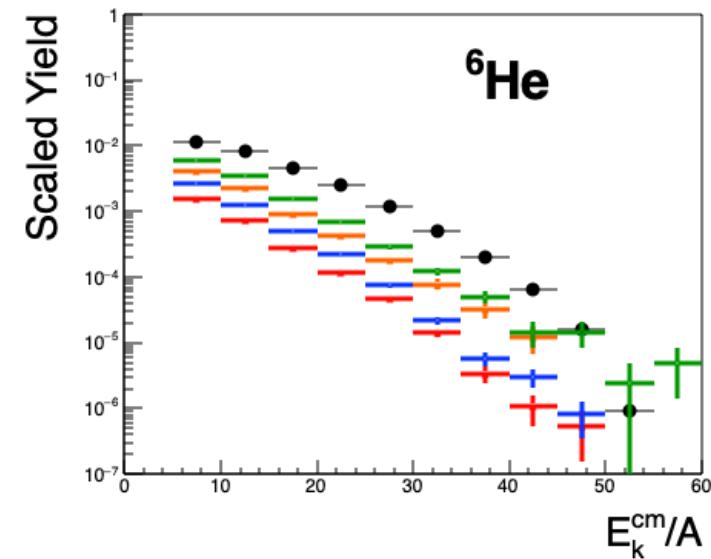
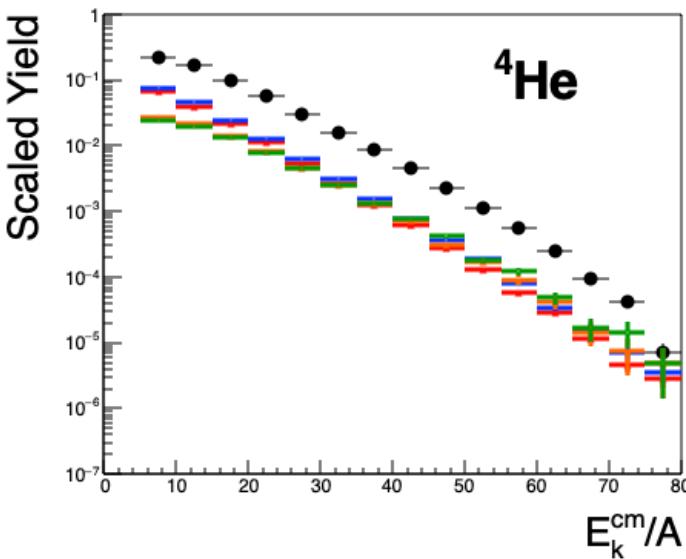
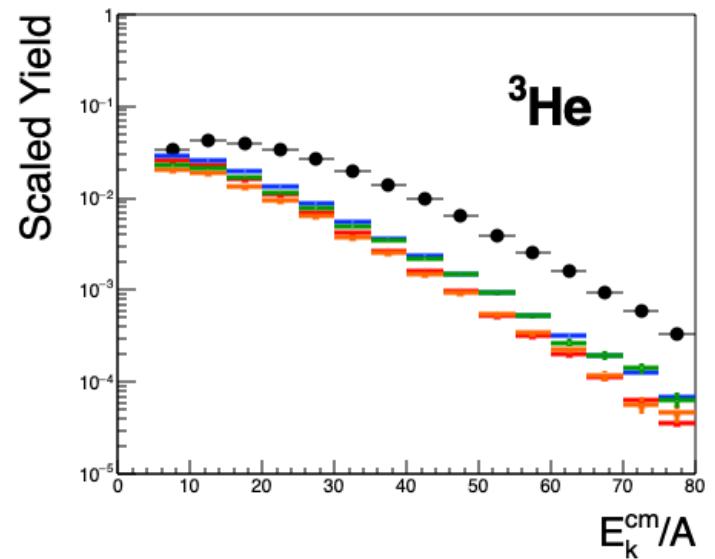
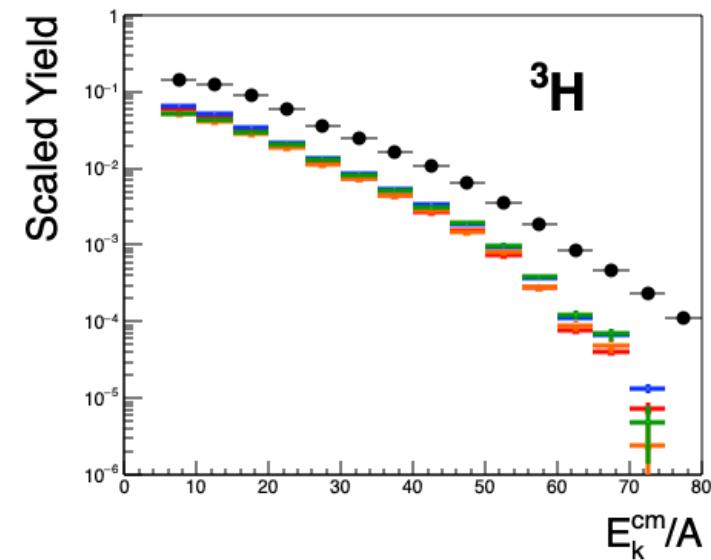
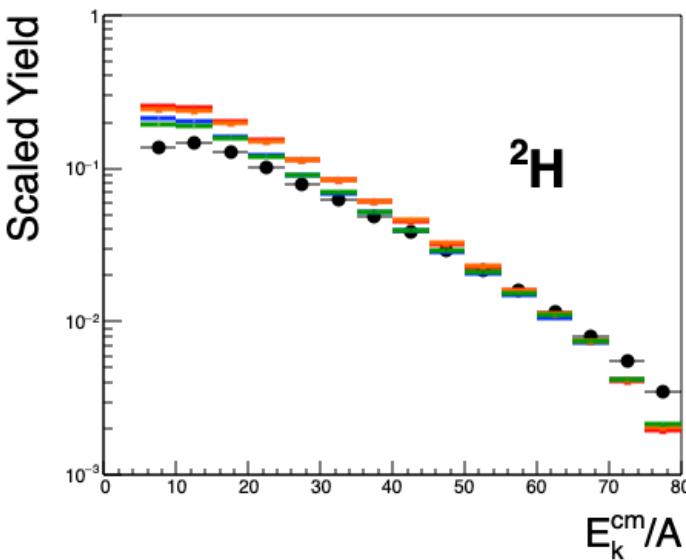
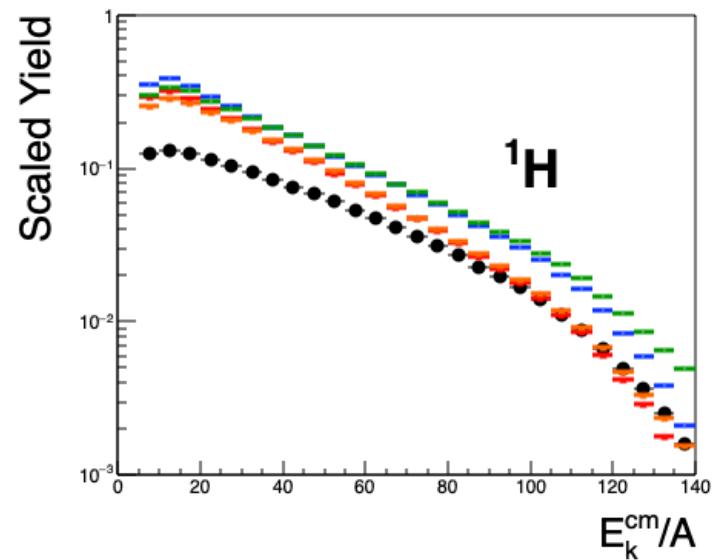


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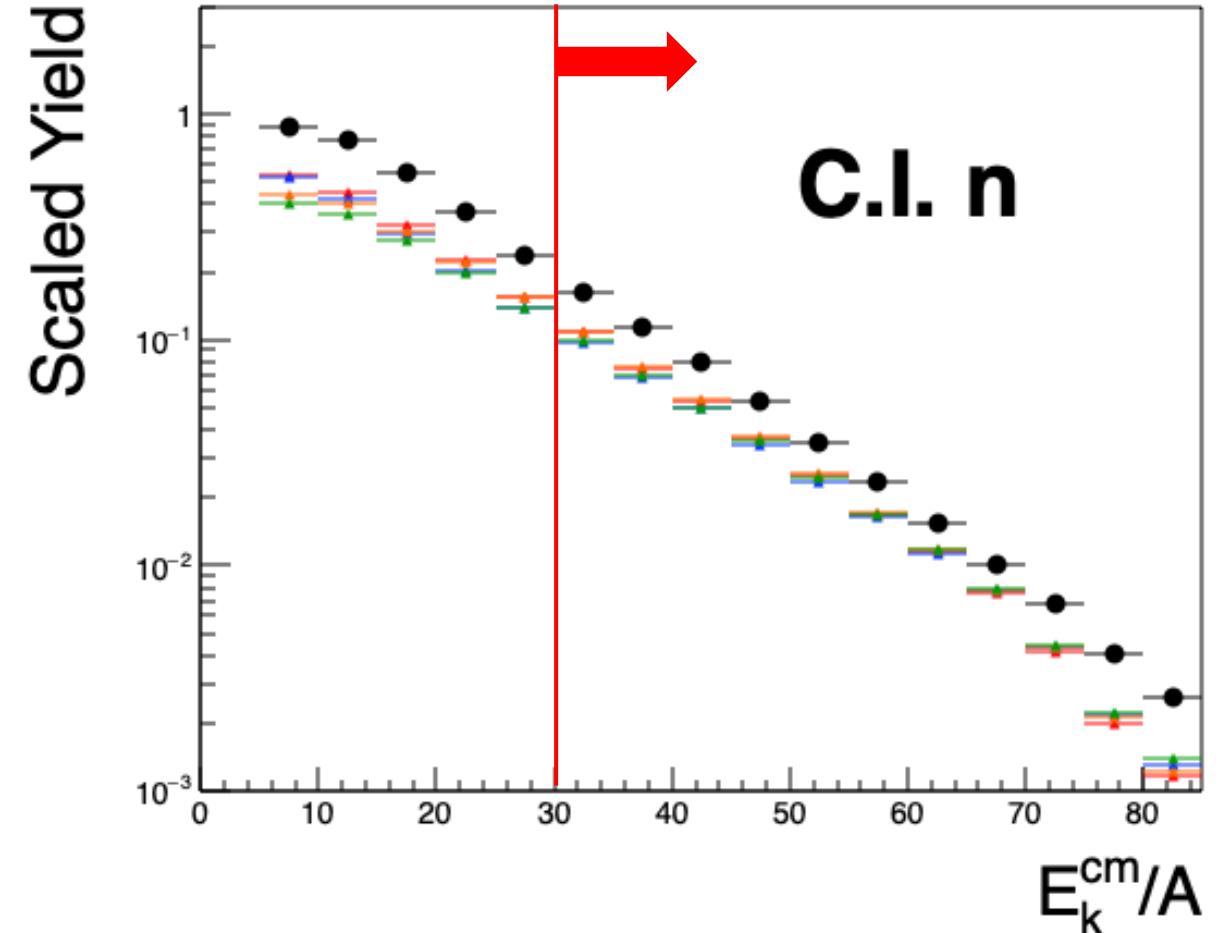


● EXP — SLy4 — SkM*
— SLy4 + GEMINI++ — SkM* + GEMINI++

129Xe+124Sn

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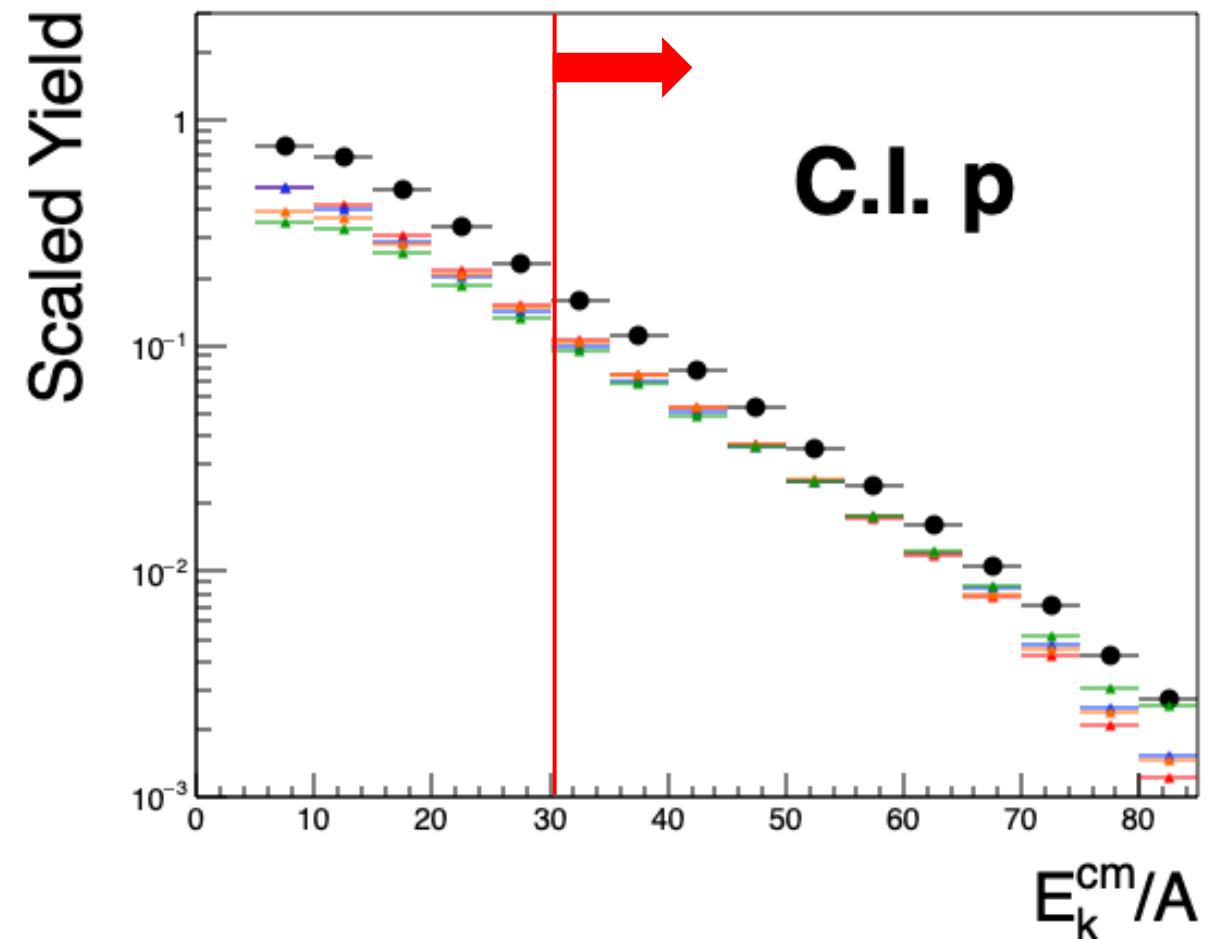


129Xe+124Sn

● EXP

— SLy4 + GEMINI++

— SkM* + GEMINI++



C.I. using $2 \leq A \leq 4$